

**Photovoltaic Quantum Well
Infrared Photodetectors**

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Quantum well infrared photodetectors are a promising new approach to long-wavelength infrared detector arrays. Both single-well photovoltaic and multiple-well photoconductive devices have been demonstrated. I will discuss noise considerations as they apply to photovoltaic devices, grating coupling of the infrared light into QWIPs, and recently demonstrated electrically tunable detectors. The use of "light trapping" to enhance the quantum efficiency and reduce cross-talk in an array will also be addressed.

LONG WAVELENGTH QUANTUM WELL DETECTORS

(Quantum Well Infrared Photodetectors)
= QWIP's

S. A. Lyon

Electrical Engineering, Princeton

Theory &
Measurements

Keith Goossen – now
at AT&T Bell Labs

Sanjay Parihar
– Princeton

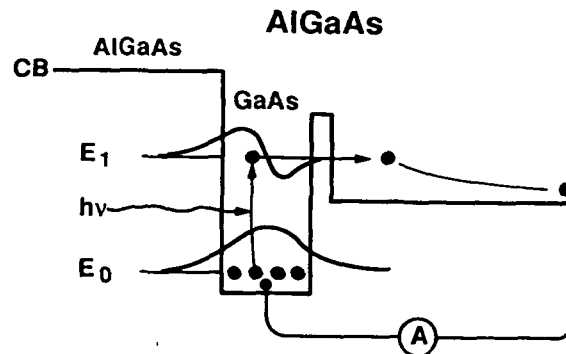
Materials

Kambiz Alavi (Siemens)
– now at U. of Texas

Mike Santos &
Mansour Shayegan
– Princeton

- I. Background on IR detectors and Quantum Wells
- II. Single Well detectors
- III. Grating enhanced detectors
- VI. Voltage Tunable Detectors
- V. Summary

QUANTUM WELL DETECTOR

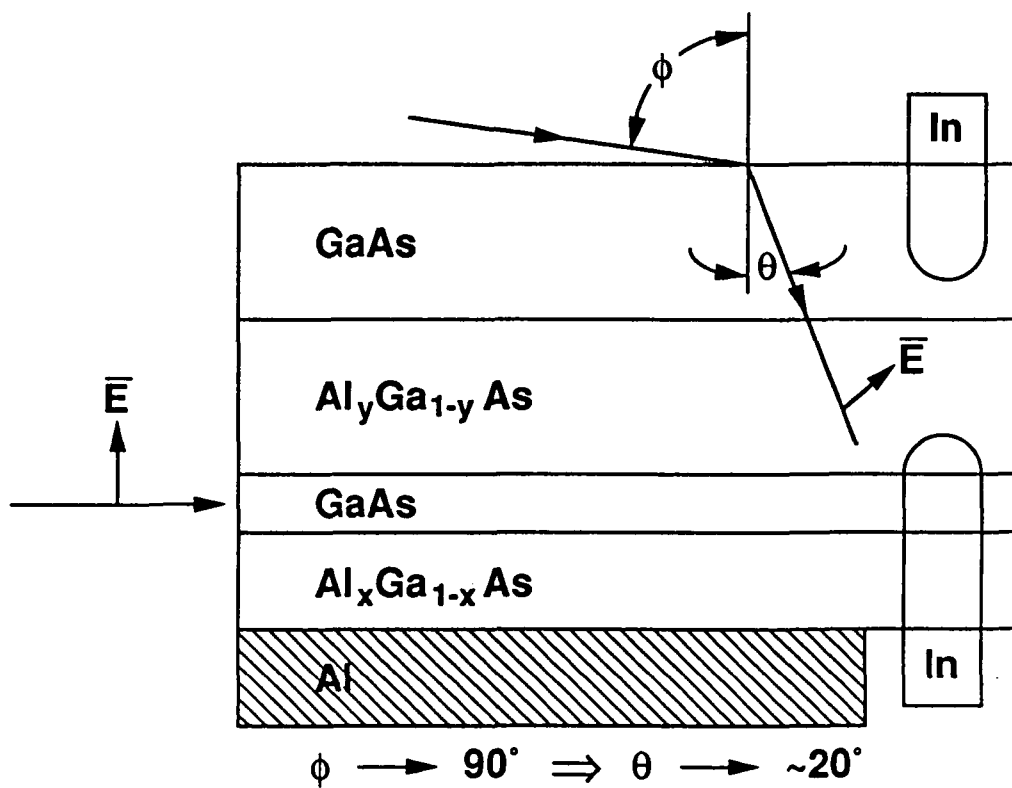
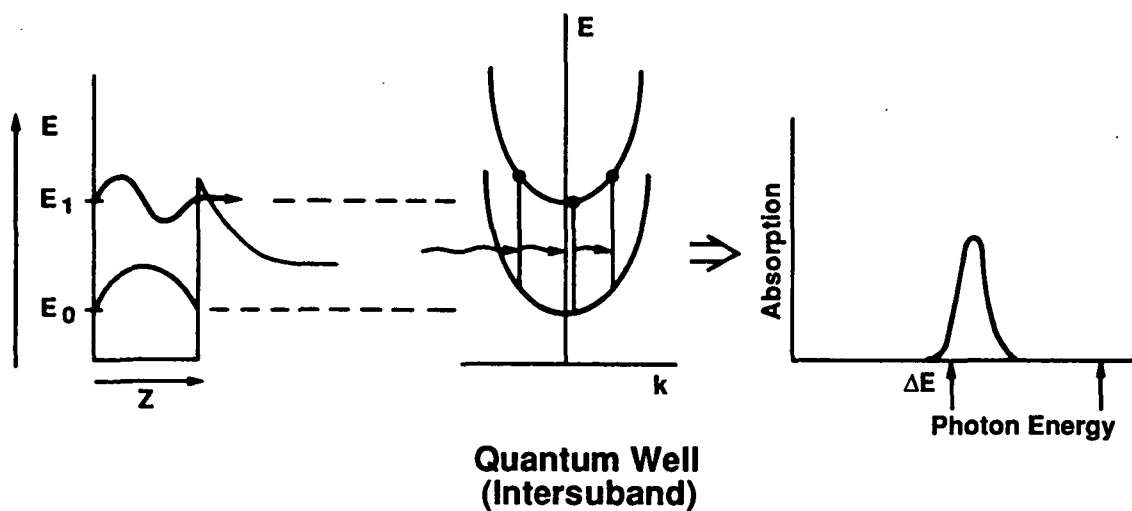
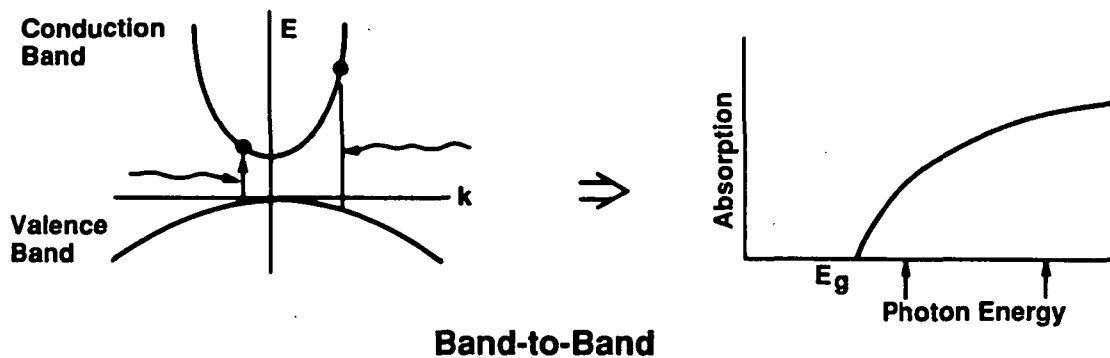


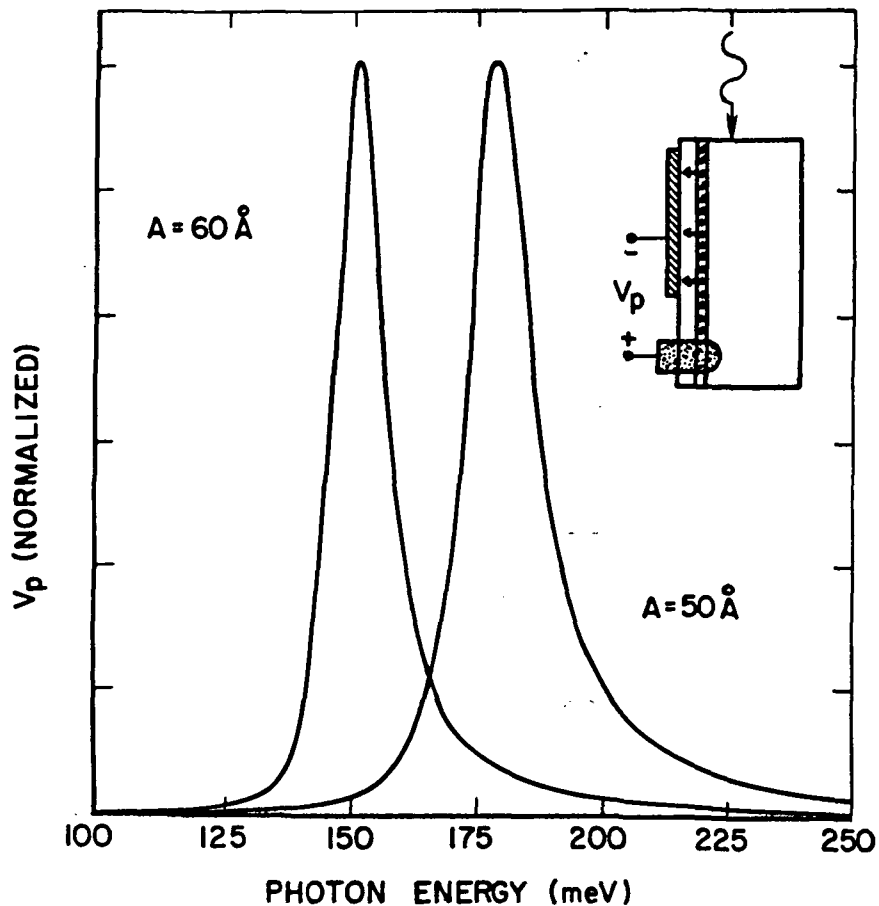
Advantages:

- Easily change structure for different wavelengths
- Long wavelength sensitivity with simple materials
- Voltage tunable

Problems:

- Not many electrons \Rightarrow low quantum efficiency
- Short relaxation time (intersubband scattering)
 \Rightarrow high dark current

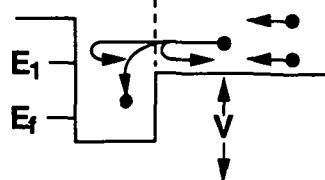




DARK CURRENT (IDEAL)

Use Richardson -Dushman approach

1. Find the rate of electron capture by the well in thermal equilibrium



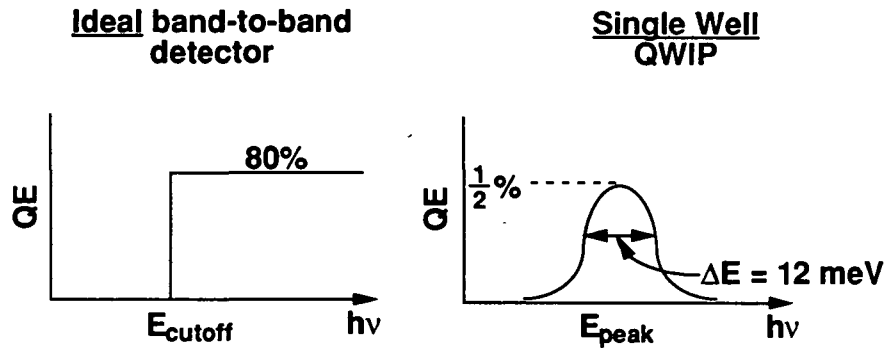
2. This will also be the emission rate \Rightarrow dark current under small biases
 3. Assume relaxation time of 1 ps
- * Unlike usual assumption for a metal:
- Capture probability < 1
 - Capture probability depends on energy

Dark current $\sim 100\times$ less for QW than metal

Work function $\sim (E_1 - E_f)$ not $(V - E_f)$

STARING ARRAY COMPARISON

50 X 50 μm Pixel F2 Optics
 30 ms frame (integration) time
 $T_{\text{Detector}} = 80\text{K}$ $T_{\text{Background}} = 300\text{K}$
 10 μm



	τ - Auger Limited (Parameters for HgCdTe)	τ - Intersubband Scattering lps
Signal ($\text{e}^-/\text{cm}^2\text{-sec}$)	3×10^{16}	3×10^{13}
Dark Current ($\text{e}^-/\text{cm}^2\text{-sec}$)	1.0×10^{14}	1.5×10^{15}
Signal/Noise	1.2×10^5	1×10^3 (NETD = 0.1K)

IMPROVING QUANTUM EFFICIENCY

1. Multiple Wells

Absorption \propto # wells

Dark Noise $\propto \sqrt{\text{\# wells}}$ (ideal)

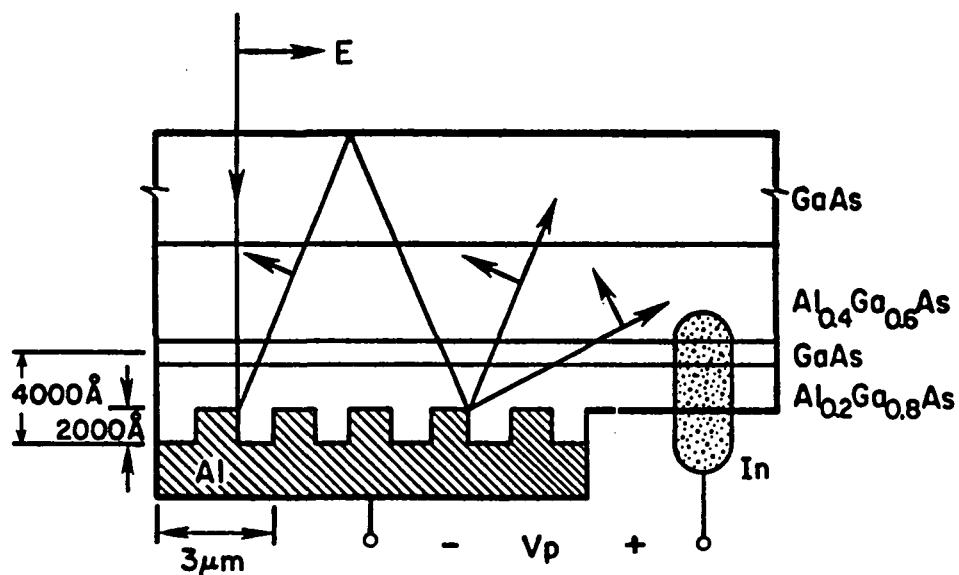
$\therefore D^* \propto \sqrt{\text{\# wells}}$

2. "Light Trapping"

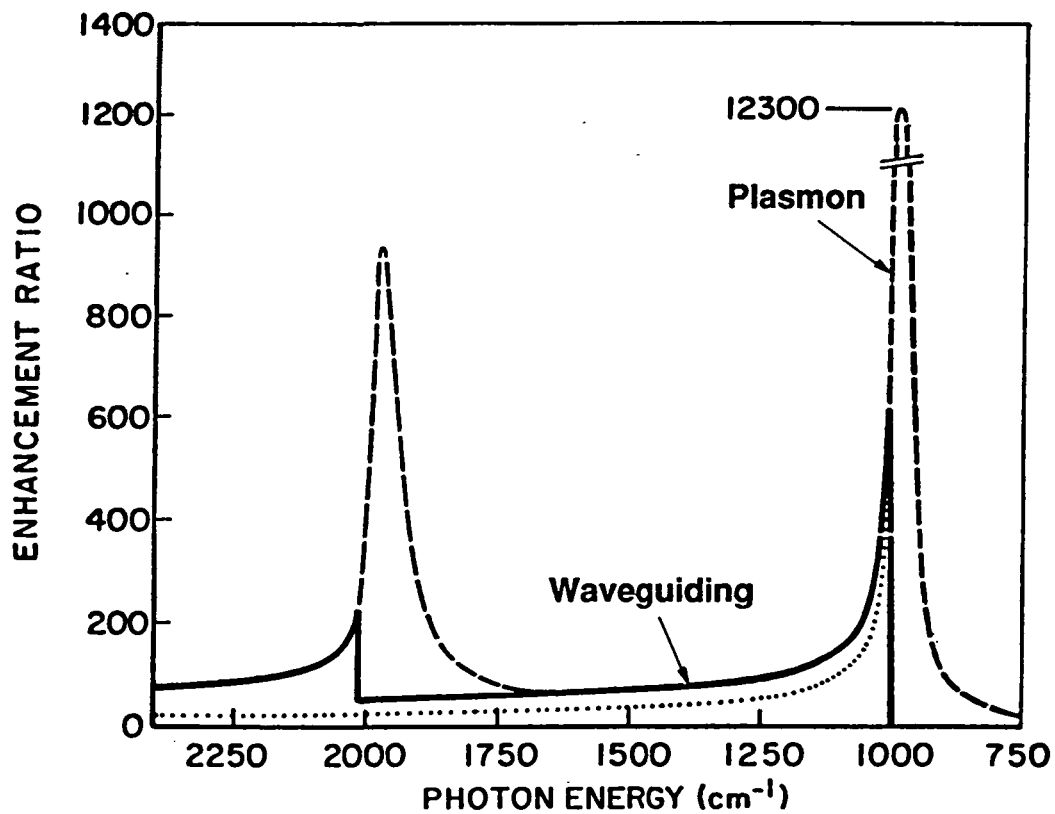
Surface Plasmons on gratings

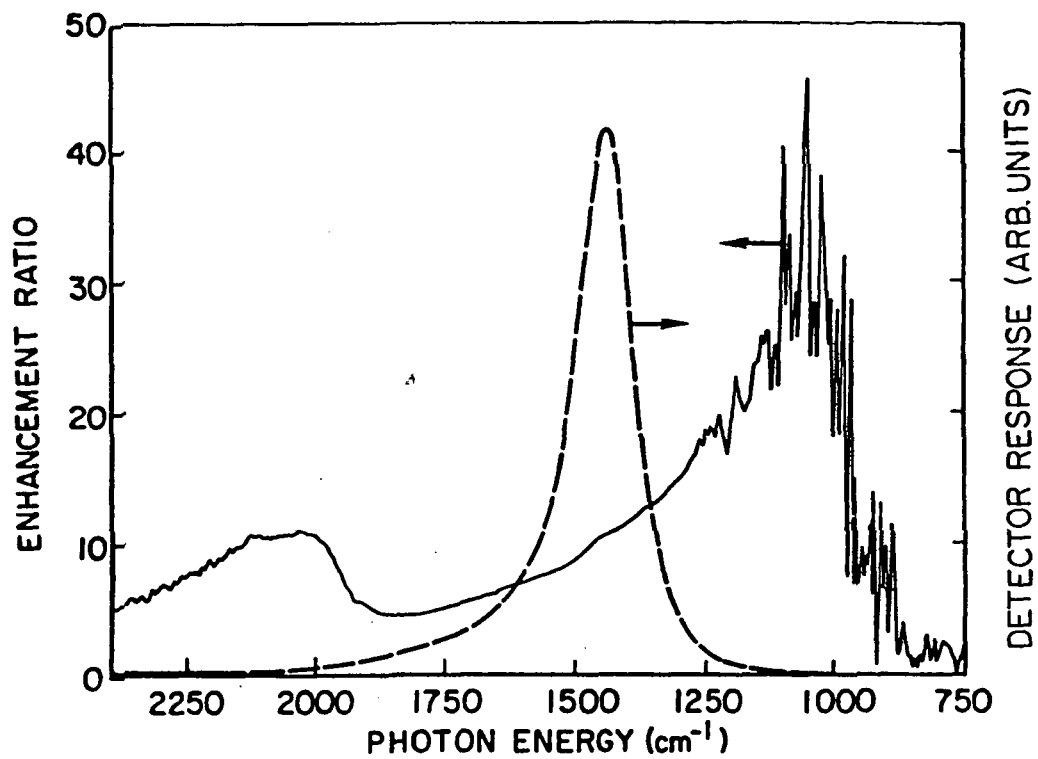
Waveguiding in GaAs

WAVEGUIDING

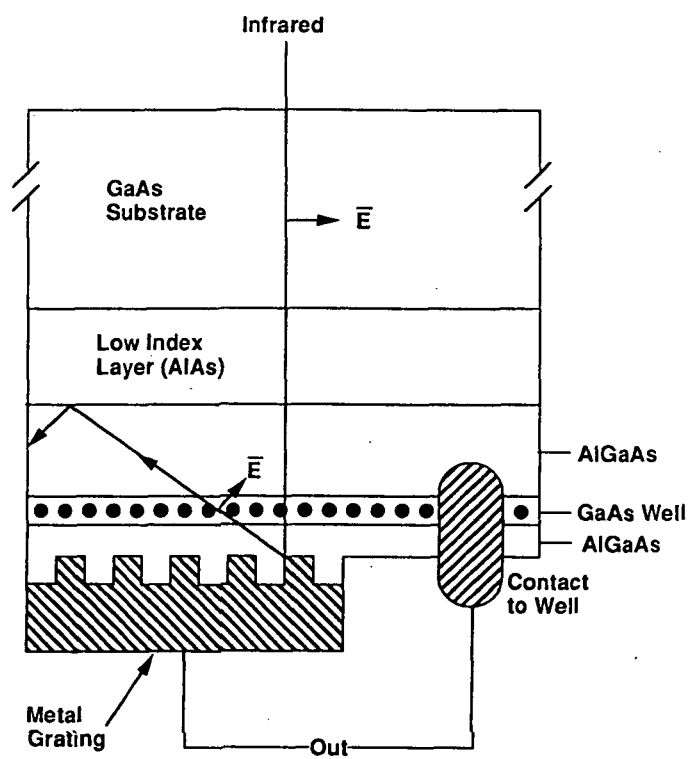


K. W. Goossen, S. A. Lyon, and K. Alavi
Appl. Phys. Lett. **53**, 1027 (1988)

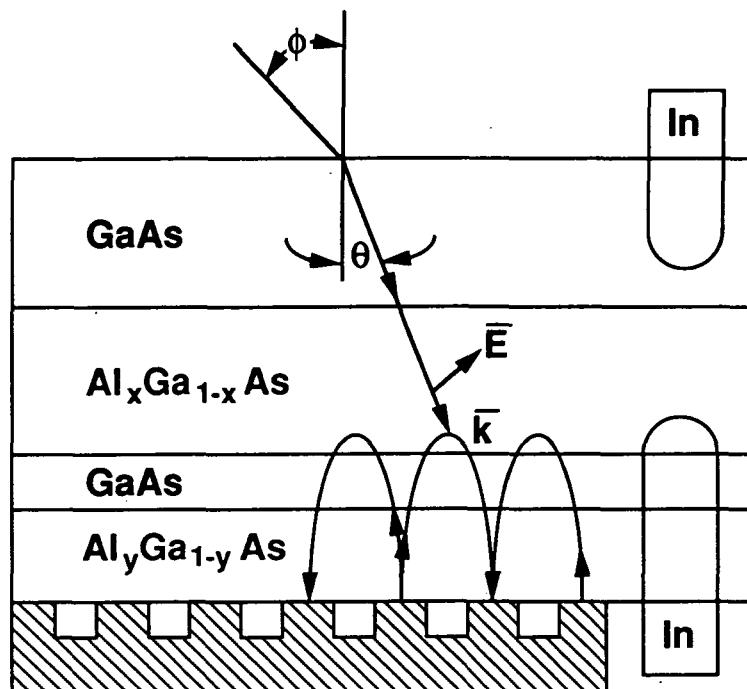




WAVEGUIDING

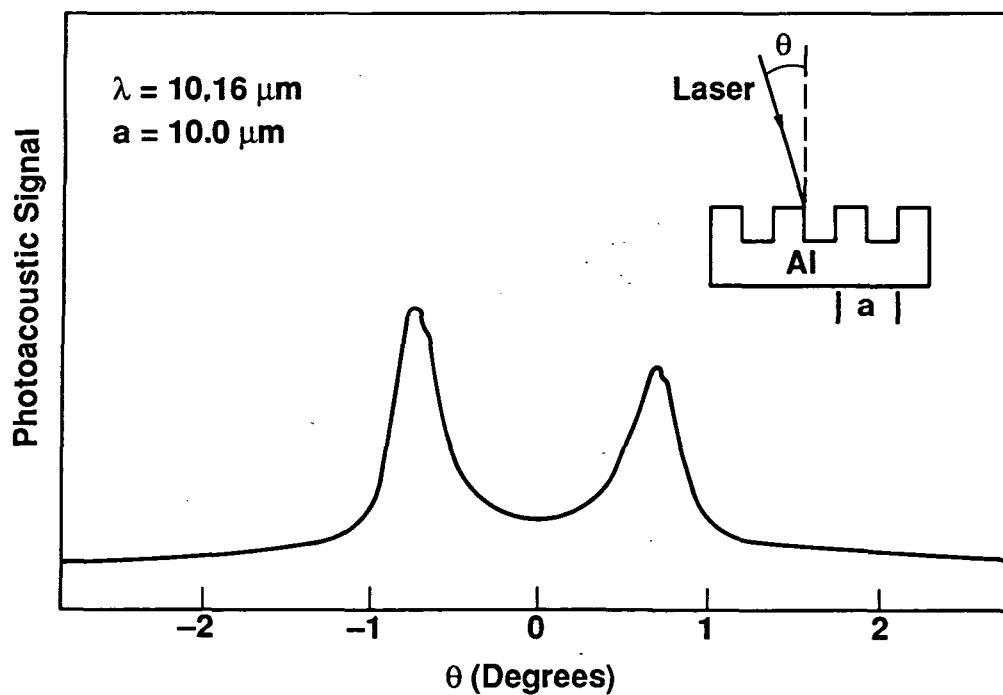


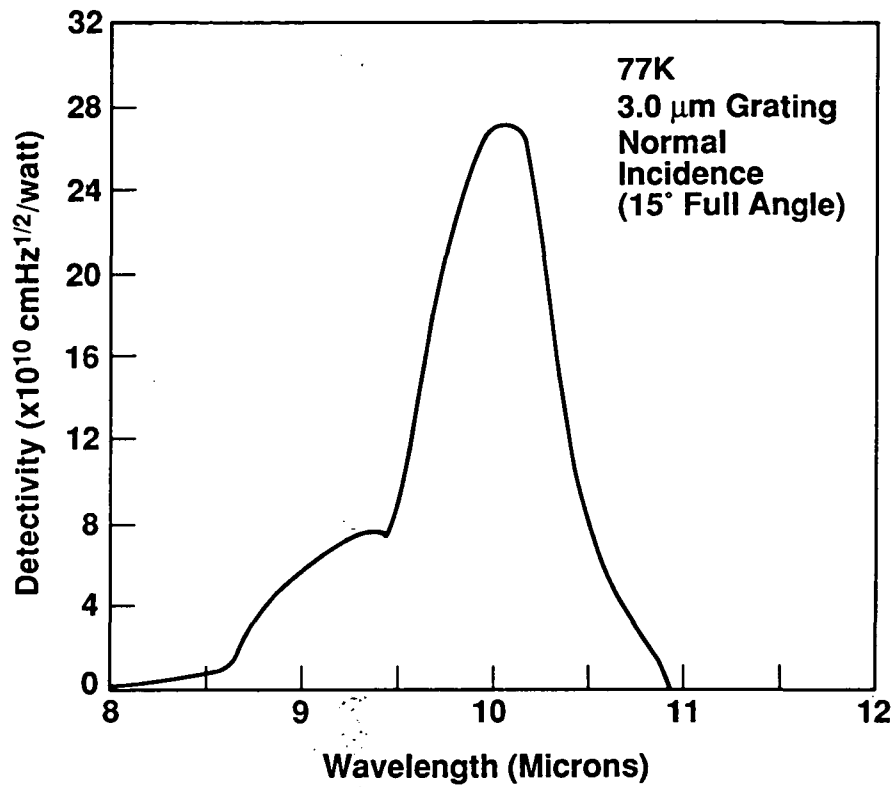
PLASMON ENHANCEMENT



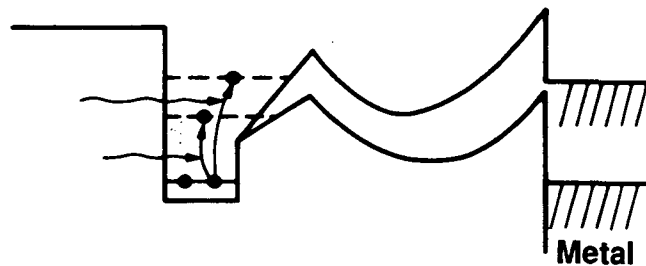
K. W. Goossen and S. A. Lyon
 Appl. Phys. Lett. 47, 1257 (1985)

SURFACE PLASMON EXCITATION

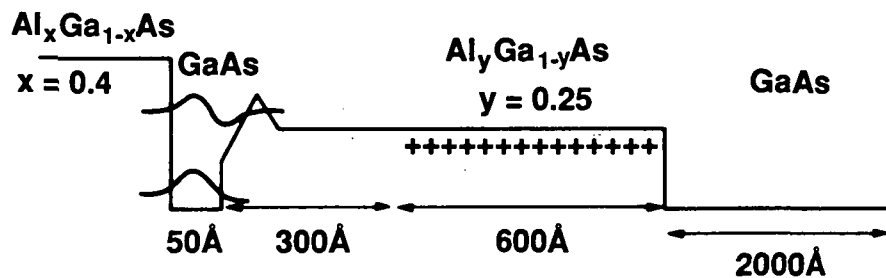




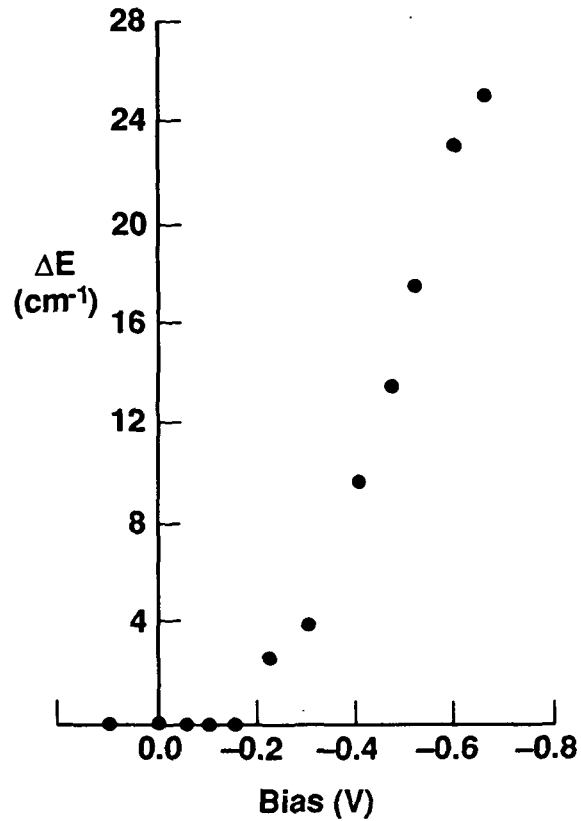
TUNABLE DETECTOR



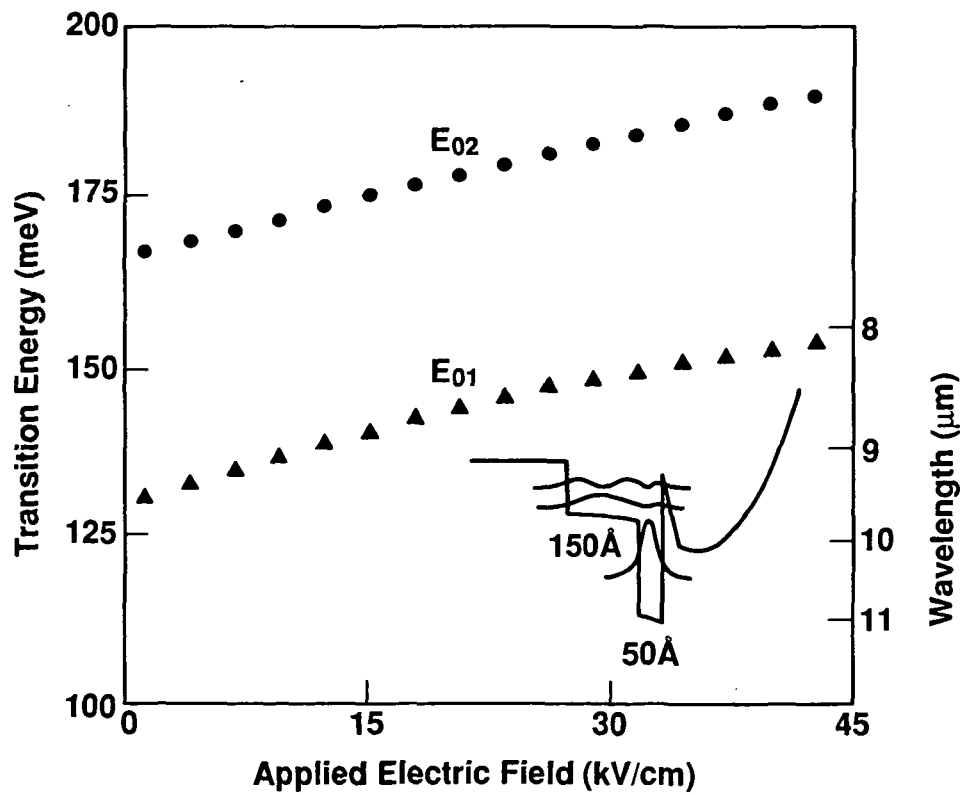
Actual Sample
(Santos & Shayegan)



ENERGY SHIFT VS BIAS FOR GRADED WELL



S. A. Lyon, MSS-4 (July 1989).
 S. R. Parihar, S. A. Lyon, M. Santos and M. Shayegan
 Appl. Phys. Lett. 55, 2417 (1989)



CONCLUSIONS

1. Single-well QWIP's work as expected
 - Quantum Efficiency ($\sim 1/2 - 1\%$)
 - Dark Current
2. Incorporation of a diffraction grating allows operation at normal incidence
3. As a single element detector QWIP's cannot compete with photoconductive HgCdTe
4. Single-well QWIP's can compete in a staring array
5. Waveguiding shows promise for enhancing quantum efficiency without the noise penalty of multiple wells
6. Voltage tunable detectors have been demonstrated
 - Similar structures expected to show large optical nonlinearities